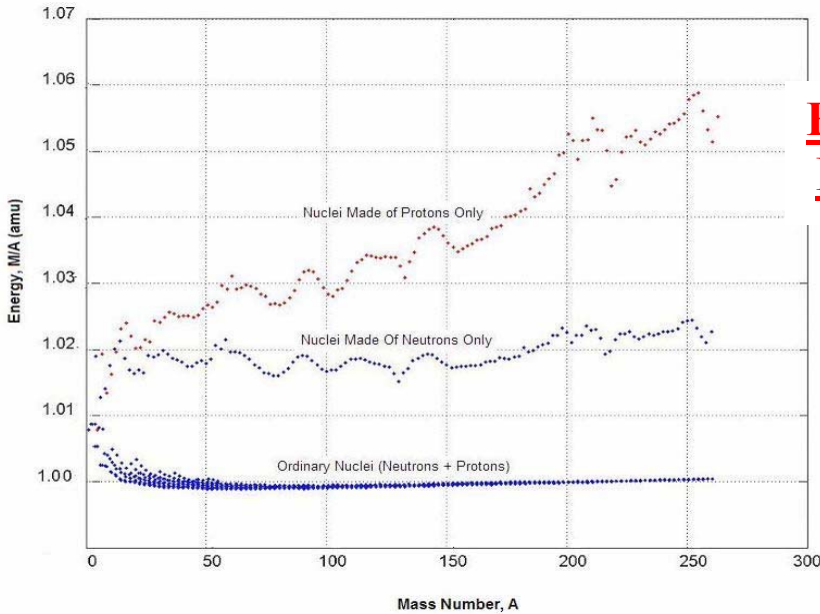


Ordinary Nuclear Matter and Nuclear Matter at $Z/A = 0$ and $Z/A = 1.0$



Energy Levels

Neutron emission from a neutron star generates ≈ 2 times the energy released in ${}^1\text{H} \rightarrow {}^4\text{He} \rightarrow {}^{56}\text{Fe}$

Sources of Solar Energy (SE) [19, 20]:

- Neutron emission from the solar core ($>57\%$ SE)
 $\langle {}^1_0\text{n} \rangle \rightarrow {}^1_0\text{n} + \sim 10\text{-}22 \text{ MeV}$
- Neutron decay or capture ($<5\%$ SE)
 ${}^1_0\text{n} \rightarrow {}^1_1\text{H}^+ + e^- + \text{anti-}\nu + 0.782 \text{ MeV}$
- Fusion and upward migration of H^+ ($<38\%$ SE)
 $4 {}^1_1\text{H}^+ + 2 e^- \rightarrow {}^4_2\text{He}^{++} + 2 \nu + 27 \text{ MeV}$
- Escape of excess H^+ in the solar wind (100% SW)
 Each year $3 \times 10^{43} \text{ H}^+ \rightarrow$ Depart in the solar wind

Neutron-emission releases **1.1% - 2.4%** of the nuclear rest mass as energy. Hydrogen-fusion releases **0.7%** if the end product is helium, and fission releases **0.1%** of the rest mass as energy [20].