The Record in the Solar Wind

Light (L) isotopes are enriched relative to heavy (H) isotopes in the solar wind by 9-stages of mass fractionation, $f = (H/L)^{4.56}$ [7].

When elemental abundances in the photosphere are corrected for this mass separation, the most abundant elements inside the Sun are Fe, Ni, O, Si, S, Mg and Ca - the elements Harkins [8] found to comprise $\approx 99\%$ of ordinary meteorites!